

3. Kinetic fluctuation theory

3.1 Background

kinetic scale, test function ϕ , empirical measure on $T_1 = \Lambda \times \mathbb{R}^2$, hard spheres diameter ϵa

$$n^\epsilon(\phi, t) = \epsilon^2 \sum_{j=1}^N \phi(q_j(t), v_j(t))$$

$$\mathbb{E}(N) = \epsilon^{-2}$$

small ϵ

$$n^\epsilon(\phi, t) \cong \int dx dv \phi(x, v) \underbrace{f(x, v, t)}_{\text{solution to B-equation}} + \epsilon \underbrace{\zeta(\phi, t)}_{\text{fluctuations}}$$

solution to B-equation

fluctuations

why density fluctuations
light scattering
equilibrium, non-equilibrium
 $T_1 \ll \ll T_2$

How to guess the structure of $\zeta(\phi, t)$?

guess of transport equation is easy, fluctuations, less clear

one expects: (i) CLT Gaussian infinite-dimensional

(ii) Markov property in t

REMINDER: \mathbb{R}^n Gauss-Markov

$$d\zeta_t = A \zeta_t dt + B dW_t$$

stationary

$\zeta_t \in \mathbb{R}^n$, $A \in M_n$, $B \in M_n$, $\{W_{t,s} \mid s \in [0, t]\}$ i.i.d. standard BM, "white noise"

• stationary measure,

$$\zeta_t(\phi) = \zeta_0(e^{A^T t} \phi) + \int_0^t ds \langle e^{A^T(t-s)} \phi, B dW_s \rangle$$

$$\mathbb{E}(\zeta_t(\phi) \zeta_t(\psi)) = \mathbb{E}(\zeta_0(e^{A^T t} \phi) \zeta_0(e^{A^T t} \psi)) + \int_0^t ds \langle e^{A^T(t-s)} \phi, B B^T e^{A^T(t-s)} \psi \rangle$$

$$\|e^{At}\| \leq e^{-\gamma t}, \gamma > 0, \text{ strict contraction}$$

Then

$$\lim_{t \rightarrow \infty} E(\zeta_t(\phi) \zeta_t(\psi)) = \int_0^\infty dt \langle \phi, e^{At} B B^T e^{A^T t} \psi \rangle = \langle \phi, C \psi \rangle$$

↑
covariance of Gaussian invariant measure

$$\Leftrightarrow -(AC + CA^T) = BB^T$$

In our application

$$A = R - D \quad \text{and}$$

$$RC = -CR^T \quad \text{reversible part}$$

$$DC = CD^T \quad \text{dissipative part} \quad \Rightarrow \quad 2DC = BB^T$$

Onsager (1931) Nobel prize not his words

A = linearized macroscopic equation, hence A(t)

first step stationary case, thermal equilibrium, macroscopic equation trivializes, but fluctuations do not

\Rightarrow C equilibrium, A linearization \rightarrow noise strength B.

apply to low density gas $\Lambda = \mathbb{R}^3$

thermal equilibrium, Gibbs measure \cong Poisson intensity $\epsilon^{-2} \rho \quad \rho > 0$

fluctuation field

$$\zeta^\epsilon(\phi, t) = \epsilon^{-1} \left(\epsilon^2 \sum_j \phi(q_j(t), v_j(t)) - \int dq dv \rho h_p(v) \cdot \phi(q, v) \right)$$

$\zeta^\epsilon(\phi, t)$ is space-time stationary

(a little bit formal)

existence of infinite volume Gibbs measure, o.k. at low density
 existence of infinite volume dynamics in \mathbb{R}^0 ok, \mathbb{R}^2 ok, higher dimensions difficult
 Prussner, 1976 equilibrium dynamics L^2 -theory
 Dobrushin, Sznur, Suhr, 1989,

Poisson

$$\lim_{\epsilon \rightarrow 0} \mathbb{E}(\xi^\epsilon(\phi, 0) \xi^\epsilon(\psi, 0)) = \int dx dv \rho h_p(v) \phi(x, v) \psi(x, v) = \langle \phi, C \psi \rangle$$

linearization

$$\rho h_p(v) + \phi(x, v)$$

$$\begin{aligned} \partial_t (\rho h_p + \phi) &= -v \cdot \nabla_x (\rho h_p + \phi) + Q(\rho h_p + \phi, \rho h_p + \phi) \stackrel{\text{linear order}}{\approx} \\ &= -v \cdot \nabla_x \phi + a^2 \int dv_1 \int d\hat{\omega} (\hat{\omega} \cdot (v - v_1))_+ \rho [h_p(v_1) \phi(v) + h_p(v) \phi(v_1) \\ &\quad - h_p(v_1) \phi(v) - h_p(v) \phi(v_1)] \\ &= R \phi - D \phi \end{aligned}$$

check

$$RC = -CR^T, \quad DC = CD^T$$

and

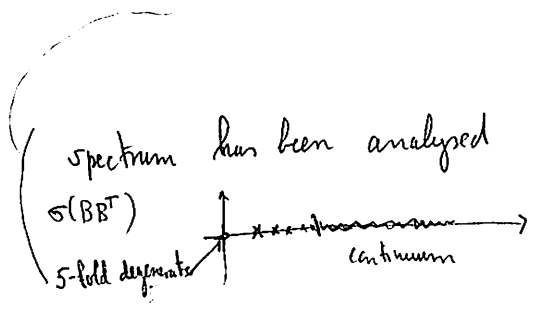
$$\begin{aligned} \langle \psi, BB^T \phi \rangle &= \frac{1}{2} \int dx \int dv_1 dv_2 \int d\hat{\omega} (\hat{\omega} \cdot (v_1 - v_2))_+ \rho h_p(v_1) \rho h_p(v_2) \\ &\quad [\psi(x, v_1') + \psi(x, v_2') - \psi(x, v_1) - \psi(x, v_2)] \\ &\quad [\phi(x, v_1') + \phi(x, v_2') - \phi(x, v_1) - \phi(x, v_2)] \end{aligned}$$

- noise is δ -correlated in time (by definition)
- δ -correlated in space
- correlated in velocities

BB^T has five eigenvalues 0 in $L^2(\mathbb{R}^3, dv)$

$$1, \bar{v}, \frac{1}{2}v^2$$

these are the conserved fields
their fluctuations are surface-like, much smaller



$$\mathcal{H} = L^2(\mathbb{R}^3, dv)$$

// structure is $BB^T = K + V$
compact multiplication

H. Grad
Ehrenfest, Illner, Pulvirenti

I do not know a simple formula for B by itself

time-dependent profile $f(x, v, t)$

linearization as $f(x, v, t) + \phi(x, v, t)$

noise-strength is time-dependent $B(t) B(t)^T$

local Poisson, natural guess

$\rho h_p(v_1) \rho h_p(v_2)$ is replaced by $f(x, v_1, t) f(x, v_2, t)$.

In our context Onsager's insight can be partially proved.

3.2 Time-dependent covariance in equilibrium

see van Beijeren, Lanford, Lebowitz, HS 1980 finite kinetic time

T. Bodineau, J. Gallagher, L. Saint-Raymond, 2016 dimension $d=2$
all kinetic times (even longer)

Theorem hard spheres, either $d \geq 2, |L| \leq \frac{1}{\epsilon} t_0$, or $d=2$; all t

$\phi, \psi \in L^2(\mathbb{R}^3, \rho h_p(v) dv)$. Then

$$\lim_{\epsilon \rightarrow 0} \mathbb{E}(\xi^\epsilon(\phi, t) \xi^\epsilon(\psi, 0)) = \langle \phi, e^{Rt - D|L|} \psi \rangle_{L^2}$$

Proof:

$$\mathbb{E}(\xi^\epsilon(\phi, t) \xi^\epsilon(\psi, 0)) = \mathbb{E} \left[\epsilon^2 \sum_j \phi(q_j(t), v_j(t)) \left(\sum_j \psi(q_j, v_j) - \mathbb{E} \left(\sum_j \psi(q_j, v_j) \right) \right) \right]$$

↑ defines a signed measure

with rescaled correlation functions

average first correlation function

$$r_n^\epsilon(x_1, \dots, x_n) = \sum_{i=1}^n \varphi(x_i) T_{eq, n}^\epsilon(x_{11}, \dots, x_n) \quad \text{decay as } e^{-|q|/\epsilon} \text{ or range of } V$$

$$+ \epsilon^{-2} \int dx_{n+1} \varphi(x_{n+1}) \left(T_{eq, n+1}^\epsilon(x_{11}, \dots, x_{n+1}) - T_{eq, n}^\epsilon(x_{11}, \dots, x_n) T_{eq, 1}^\epsilon(x_{n+1}) \right)$$

we write $T_{eq, n}^\epsilon(x_{11}, \dots, x_n) = (q_{11}, \dots, q_n) \prod_{j=1}^n h_p(v_j)$

From the low density expansion (see book by Ruelle)

$$\sup_{\vec{q} \in \mathbb{R}^3} \epsilon^{-2} \int_{\mathbb{R}^3} dq |T_{eq, n+1}^\epsilon(\vec{q}, q) - T_{eq, n}^\epsilon(\vec{q}) T_{q, 1}^\epsilon(q)| \leq C \epsilon$$

uniformly in the volume

→ uniform bound for $0 \leq t \leq c^{-1} L_0$

→ term by term convergence

limit correlation functions satisfy Boltzmann hierarchy with initial conditions

$$r_n(x_1, \dots, x_n) = \sum_{j=1}^n \varphi(x_j) \prod_{i=1}^n h_p(v_i)$$

$$\cong \left(e^{\sum_{j=1}^n \varphi(x_j)} - 1 \right) \prod_{i=1}^n h_p(v_i) \quad \text{to linear order}$$

→

$$r_n(x_1, \dots, x_n, t) = \sum_{j=1}^n (e^{(A-D)t} \varphi)(x_j) \prod_{i=1}^n h_p(v_i)$$

$n=1$ yields the desired limit. □

- A similar argument confirms also the time-dependent covariance.
- stochastic process, open problem
- dynamical correlations on the kinetic scale → interparticle distance

3.3 Beyond the kinetic time scale

new theme, but with preexamples, e.g. Erdős, Yau on quantum particle in a random potential

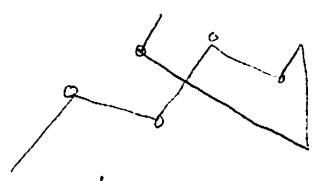
Simple Example random Lorentz gas
 \mathbb{R}^d Poisson point process, intensity $\varepsilon^{-(d-1)}$ $\{q_j\}$

$\hat{=}$ central scatterers, diameter of scatterer is ε
 mechanical motion in this random potential



$$\frac{d^2}{dt^2} q^\varepsilon(t) = \sum_j F_\varepsilon(q^\varepsilon(t) - q_j), \quad q^\varepsilon(0) = 0 \quad F_\varepsilon = -\nabla V_\varepsilon, \quad V_\varepsilon(x) = V(x/\varepsilon)$$

V radial
here hard core



$$q^\varepsilon(t) = \int_0^t v^\varepsilon(s) ds \quad v = \dot{q} \quad \text{specific observable}$$

$\lim_{\varepsilon \rightarrow 0} 0 \leq t \leq T$ as a stochastic process
 $\lim_{\varepsilon \rightarrow 0} v^\varepsilon(t) = v(t)$

$v(t)$ is a random jump process, $|v(t)| = |v(0)| = 1$, e.g. $d=3$
 $v(t)$ exponential waiting, uniformization of velocities over the sphere.

holds for all T .

Q: Over which time ^{span} remains $q^\varepsilon(t)$ close to $q(t) = \int_0^t v(s) ds$
 ↑ mechanical ↑ random flight

more sharply: take $\alpha > 0$, consider

$$\frac{1}{\sqrt{\varepsilon^{-\alpha} t}} q^\varepsilon(\varepsilon^{-\alpha} t)$$

If this stays close, then $q^\varepsilon(t)$ is already Brownian motion $B(t)$.

Problem: For which α

$$\lim_{\varepsilon \rightarrow 0} \frac{1}{\sqrt{\varepsilon^{-\alpha} t}} q^\varepsilon(\varepsilon^{-\alpha} t) = B(t)$$

← from Markov limit

recent result of B. Toth claims $0 < \kappa < 1$, presumably $0 < \kappa < 2$

BGS 2016 prove convergence to linearized Boltzmann for

t fixed, $N = \varepsilon^{-1} \nu$, $\nu \rightarrow \infty$ but $\nu \ll \sqrt{\log \log \log N}$

on that scale one observes ^{already} the hydrodynamic approximation to the linearized Boltzmann equation

$$r_1^{(N)}(x, v, 0) = h_p(v) \psi(x, v, 0)$$

mean zero, internal energy, $d=2$

with

$$\psi(x, v, 0) = \rho_0(x) + u_0(x) \cdot v + \theta_0(x) \left(\beta \frac{1}{2} v^2 - 1 \right)$$

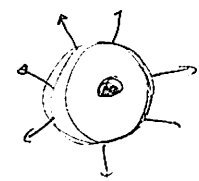
Then

$$\lim_{\substack{N \rightarrow \infty \\ \nu \rightarrow \infty}} r_1^{(N)}(x, v, t) = h_p(v) \left(\rho(x, t) + \tilde{u}(x, t) \cdot v + \theta(x, t) \left(\beta \frac{1}{2} v^2 - 1 \right) \right)$$

and ρ, u, θ solves

$$\begin{aligned} \partial_t \rho + \nabla_x \cdot \tilde{u} &= 0 \\ \partial_t \tilde{u} + \nabla_x (\rho + \theta) &= 0 \\ \partial_t \theta + \nabla_x \cdot u &= 0 \end{aligned}$$

(. sound and energy waves)



(linearized hydrodynamics).

Lecture 6, Oct 31, 10:00

nonlinear waves