

(Large scale dynamics and their fluctuation theory)

lectures in recent years on stochastic particle systems, dynamics of interfaces, KPZ related topics

other area of active interest

carrying coal to Newcastle
hot dogs to New York

• dynamics of many classical particles

with random initial conditions

dynamical system

dynamical system \Leftrightarrow probability

law of large numbers

goal is fluctuation theory

low density fluids

|| weakly nonlinear waves

|| one-dimensional fluids

Gaussian

Gaussian

nonlinear fluctuation theory

large deviations?

Large Scale Dynamics and Their Fluctuation Theorems

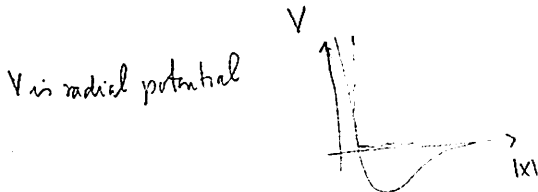
1. A puzzle from Statistical Physics

many-body mechanical system

positions $q_j \in \mathbb{R}^3$, momenta (= velocities) $p_j \in \mathbb{R}^3$, $j=1, \dots, N$, mass m

Newton 1686, Maxwell, Boltzmann \approx 1860

$$\frac{d}{dt} q_j = \frac{1}{m} p_j, \quad \frac{d}{dt} p_j = - \sum_{\substack{k=1 \\ k \neq j}}^N \nabla V(q_j - q_k), \quad j=1, \dots, N \quad (\text{Newton}) + \text{random initial condition}$$

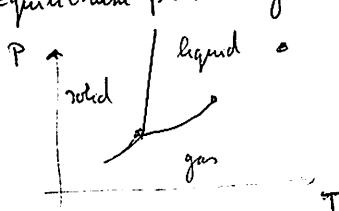


V is radial potential

model potential

"true" potential comes from quantum mechanics
"simple fluid"

equilibrium phase diagram (Gibbs measures)

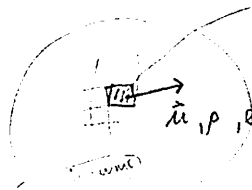


|| fluid ||

$$\frac{1}{Z} e^{-\beta \sum_{i,j} V(q_i - q_j)} \quad \text{stationary measures}$$

ρ density
 \vec{u} velocity

Euler 1757 continuum description



total energy = kinetic + internal
 $\rho(x,t)$, $x \in \mathbb{R}^3, t \geq 0$

$$\partial_t \rho + \nabla_x \cdot (\rho \vec{u}) = 0$$

$$\partial_t (m \rho u_\alpha) + \nabla_x \cdot (m \rho u_\alpha \vec{u}) + \partial_x_\alpha P(\rho, e_{int}) = 0, \quad \alpha=1,2,3 \quad (\text{Euler})$$

$$\partial_t e + \partial_x \cdot j_e = 0$$

no pressure in Newton

+ initial conditions

|| Why do particles follow the solution of the PDE?

physics: taken for granted || for study microscopic scale
fluid dynamics
Mark Kac 1959 Probability and related topics in physical sciences

hierarchical model for governing equations

fact hint 3g of water has $N = 10^{23}$

⇒ law of large numbers

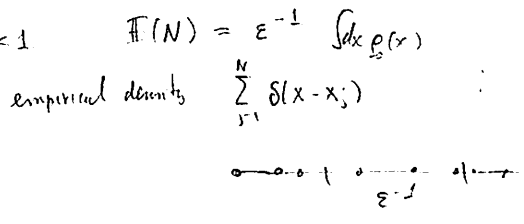
EXAMPLE (oversimplified)

only positions $x_j \in \mathbb{R}, j=1, \dots, N$

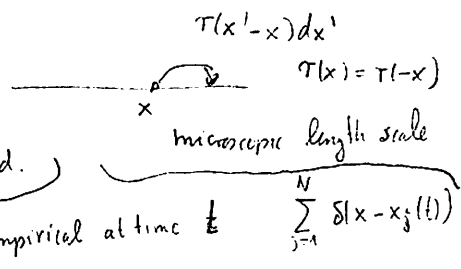
static $t=0$ $\{x_j\}$ Poisson point process intensity $\rho_0 > 0$ $\int dx \rho_0(x) < \infty, N$ random

slow variation $\rho_0(x)$ $\epsilon \ll 1$ $\mathbb{P}(N) = \epsilon^{-1} \int dx \rho_0(x)$ interparticle 1

f. test function $n^\epsilon(t) = \epsilon \sum_{j=1}^N f(\epsilon x_j) =$



$\lim_{\epsilon \rightarrow 0} n^\epsilon(t) = \int dx \rho_0(x) f(x)$ a.s.



dynamics $x_j(t)$ symmetric random walk i.i.d.

then $\lim_{\epsilon \rightarrow 0} \mathbb{E}(x_j | \epsilon^{-2}t) = \sqrt{D} \mathcal{G}(t)$ BH

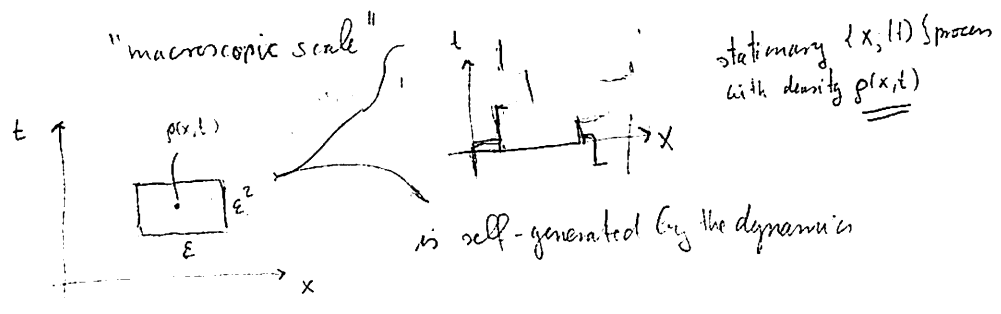
$n^\epsilon(t, t) = \epsilon \sum_{j=1}^N f(\epsilon x_j(\epsilon^{-2}t))$

$\lim_{\epsilon \rightarrow 0} n^\epsilon(t, t) = \int dx \rho(x, t) f(x)$ a.s. $\partial_t \rho = \frac{1}{2} \partial_x D \partial_x \rho$ $\rho(x, 0) = \rho_0(x)$

three lessons

⇒ law of large numbers

⇒ local stationarity (macroscopic scale)



⇒ fluctuations CLT linear fluctuation theory OU process

critics

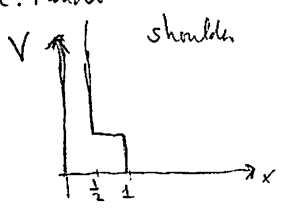
• macroscopic equation is linear, nonlinear interactions, HERE $D \rightarrow D(\rho)$
 Varadhan, Yau, ... hydrodynamic limit

• physical evolution is deterministic BUT with random initial data

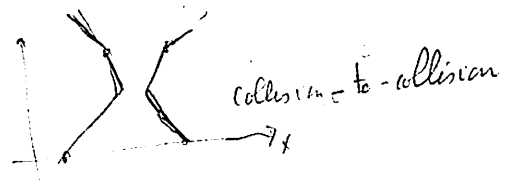
• fluid is not local Poisson, equilibration poorly understood for deterministic systems

molecular dynamics
work with C. Mendell

d=1
N=4000



density $\frac{1}{2}$



$$\frac{1}{Z} e^{-\beta \sum V(q_{3i2} - q_s)}$$

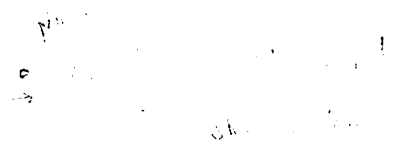
particle positions equilibrated, velocities z.i.d. not Gaussian

t = 2000 collisions per particle, study local statistics, one particle velocity distribution

z.i.d. Gaussian is observed

(3 body collisions required)
the denser the better

depends on β



... ..

currently: Euler is like Mt Everest, no trail

Olla, Varadhan, Yau 1993
still needs stochastic collisions
best available, no follow up

HERE, small parameter,

1) low density

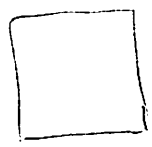
Lanford 1976
much recent work.

2) weakly nonlinear wave equation

J. Lukkarinen, HS. 2005-2009

2. Low density gas

2.1 Kinetic scaling



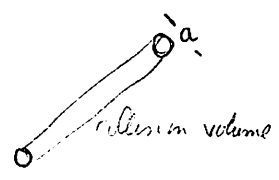
Box $\Lambda \subset \mathbb{R}^3$ kinetic scale (macroscopic)

N hard spheres diameter ϵa $\epsilon \ll 1$

$E(N) \rightarrow \infty$

• velocities are $O(1)$

• N and ϵ ? mean free path fixed



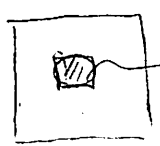
$(\epsilon a)^2 l N = |\Lambda|$

$\approx \epsilon^2 N = O(1)$

volume occupied $(\epsilon a)^3 \epsilon^{-2} = \epsilon a^3 \rightarrow 0$ as $\epsilon \rightarrow 0$

ideal gas thermodynamics

Gibbs measure: $\frac{1}{Z} \frac{1}{N!} dq_1 \dots dq_N \exp\left[-\frac{1}{2} \sum_{i+j=1}^N V((q_i - q_j)/\epsilon)\right] = \mu^\epsilon$



$B(l)$

box with a finite number of particles

$l^3 N = 1 \quad l = \epsilon^{2/3}$

$\mu^\epsilon |_{B(l)}$ $\xrightarrow{\text{rescaled}}$ Poisson process

$\delta > 0$

$N = \epsilon^{-2-\delta}$ strong interactions

$N = \epsilon^{-2+\delta}$ non-interacting limit

|| free motion / collisions on the same scale

end of lect. 1

2.2 The rate equation

one-particle phase space $\Lambda \times \mathbb{R}^3$, Boltzmann f function $f: \Lambda \times \mathbb{R}^3 \rightarrow \mathbb{R}_+$, infinite volume $\Lambda = \mathbb{R}^3$

$N \int_{\Delta} dx dv f(x,v,t) \cong$ number of particles in box $\Delta \subset \Lambda \times \mathbb{R}^3$

notation $x \in \Lambda, v \in \mathbb{R}^3$

change in time of f

1) free streaming $\partial_t f = -v \cdot \nabla_x f$